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8.3 Nuclear Instability &



PHYSICS

AQA A Level Revision Notes



A Level Physics AQA

8.3 Nuclear Instability & Radius

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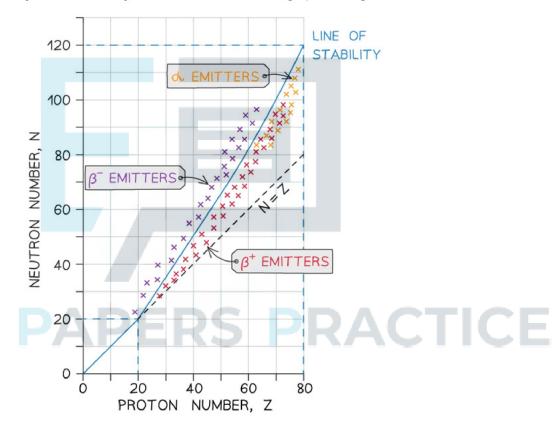
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8.3.1 Nuclear Instability

Nuclear Stability Graph

- The most common elements in the universe all tend to have values of N and Z less than 20 (plus iron which has Z = 26, N = 30)
- · Where:
 - N = number of neutrons
 - Z = number of protons / atomic number
- This is because lighter elements (with fewer protons) tend to be much **more stable** than heavier ones (with many protons)
- Nuclear stability becomes vastly clearer when viewed on a graph of N against Z



EXAM

This nuclear stability curve shows the line of stable isotopes and which unstable isotopes will emit alpha or beta particles

- · A nucleus will be unstable if it has:
 - Too many neutrons
 - Too many protons
 - o Too many nucleons ie. too heavy
 - o Too much energy
- An unstable atom wants to become neutral to become stable
- For light isotopes, Z < 20:



- All these nuclei tend to be very stable
- They follow the straight-line N = Z
- For heavy isotopes, Z > 20:
 - The neutron-proton ratio increases
 - Stable nuclei must have more neutrons than protons
- This imbalance in the neutron-proton ratio is very significant to the stability of nuclei
 - At a short range (around 1-4 fm), nucleons are bound by the **strong nuclear force**
 - Below 1 fm, the strong nuclear force is repulsive in order to prevent the nucleus from collapsing
 - At longer ranges, the electromagnetic force acts between protons, so more protons cause more instability
 - Therefore, as more protons are added to the nucleus, more neutrons are needed to add distance between protons to reduce the electrostatic repulsion
 - Also, the extra neutrons increase the amount of binding force which helps to bind the nucleons together



Alpha, Beta & Electron Capture

- The graph of N against Z is useful in determining which isotopes will decay via
 - Alpha emission
 - Beta-minus (β⁻) emission
 - Beta-plus (β+) emission
 - Electron capture

Alpha-emitters:

- Occur beneath the line of stability when Z > 60 where there are too many nucleons in the nucleus
- These nuclei have more protons than neutrons, but they are too large to be stable
- This is because the strong nuclear force between the nucleons is unable to overcome the electrostatic force of repulsion between the protons

Beta-minus (β⁻) emitters:

- Occur to the left of the stability line where the isotopes are neutron-rich compared to stable isotopes
- A neutron is converted to a proton and emits a β⁻ particle (and an anti-electron neutrino)

Beta-plus (β⁺) emitters:

- Occur to the right of the stability line where the isotopes are proton-rich compared to stable isotopes
- A proton is converted to a neutron and emits a β⁺ particle (and an electron neutrino)

Electron capture:

- When a nucleus captures one of its own orbiting electrons
- As with β⁺ decay, a proton in the nucleus is converted into a neutron, releasing a gamma-ray (and an electron neutrino)
- Hence, this also occurs to the right of the stability line where the isotopes are protonrich compared to stable isotopes



Exam Tip

To remember where the β^- and β^+ emitters are on the graph:

- Beta-minus is a negative particle where a neutron turns into a proton. Unstable atoms always want to go towards a roughly equal number of protons and neutrons
 - Therefore these emitters are on the neutron-rich side of isotopes
- Beta-plus is a **p**ositive particle where a **p**roton turns into a neutron
 - Therefore these emitters are on the proton-rich side of isotopes

The best way to remember the nuclear stability graph is to try to draw it from memory



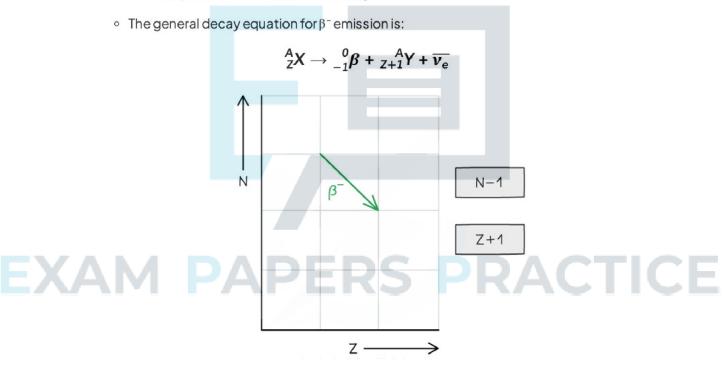
8.3.2 Decay Equations

Changes in N and Z by Radioactive Decay

• There are four reasons why a nucleus might become unstable, and these determine which decay mode will occur

· Too many neutrons

- Decays through beta-minus (β-) emission
- \circ One of the **neutrons** in the nucleus changes into a **proton** and a β^- particle (an electron) and antineutrino is released
- The nucleon number is constant
 - The neutron number (N) decreases by 1
 - The proton number (Z) increases by 1



· Too many protons

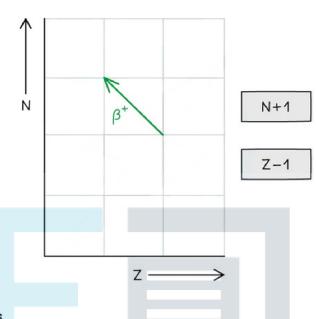
- Decays through **beta-plus** (β+) emission or electron capture
- In beta-plus decay, a **proton** changes into a **neutron** and a β^+ particle (a positron) and neutrino are released
- In electron capture, an orbiting electron is taken in by the nucleus and combined with a proton causing the formation of a neutron and neutrino
- In both types of decay, the nucleon number stays constant
 - The neutron number (N) increases by 1
 - The proton number (Z) decreases by 1
- The general decay equation for β+ emission is:



$${}^{A}_{Z}X \rightarrow {}^{0}_{+1}\beta + {}^{A}_{Z-1}Y + \nu_{e}$$

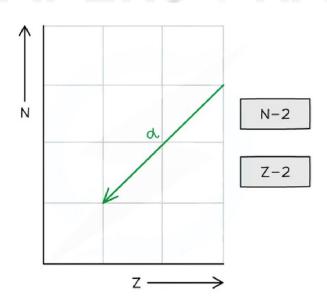
• The equation for electron capture is:

$$_{Z}^{A}X + _{-1}^{0}e \rightarrow _{Z-1}^{A}Y + \nu_{e}$$



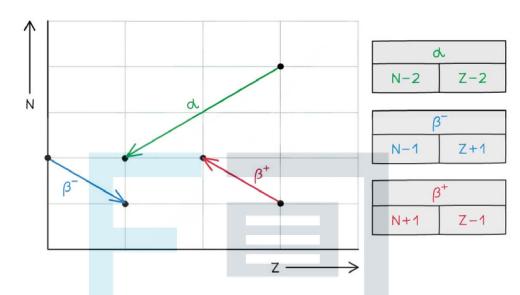
- Too many nucleons
 - Decays through alpha (α) emission
 - o An α particle is a helium nucleus
 - o The nucleon number decreases by 4 and the proton number decreases by 2
 - The neutron number (N) decreases by 2
 - The proton number (Z) decreases by 2
 - The general decay equation for α emission is:

EXAMPAP $_{zX \rightarrow \frac{4}{2}\alpha + \frac{A-4}{Z-2}Y}$ PRACTICE





- · Too much energy
 - Decays through gamma (γ) emission
 - o A gamma particle is a high-energy electromagnetic radiation
 - o This usually occurs after a different type of decay, such as alpha or beta decay
 - This is because the nucleus becomes excited and has excess energy
- In summary, alpha decay, beta decay and electron capture can be represented on an N-Z graph as follows:





Worked Example

Plutonium-239 is a radioactive isotope that contains 94 protons and emits α particles to form a radioactive isotope of uranium. This isotope of uranium emits α particles to form an isotope of thorium which is also radioactive.

- a) Write two equations to represent the decay of plutonium-239 and the subsequent decay of uranium
- b) Predict the decay mode of the thorium isotope
- c) Draw the decay chain from plutonium-239 to the daughter product of thorium decay on an N-Z graph

Part (a) Step 1: Write down the general equation of alpha decay

$${}_{7}^{A}X \rightarrow {}_{2}^{4}\alpha + {}_{7-2}^{A-4}Y$$

Step 2: Write down the decay equation of plutonium into uranium

$$^{239}_{94}$$
Pu $\rightarrow ^{4}_{2}\alpha + ^{235}_{92}$ U

Step 3: Write down the decay equation of uranium into thorium

$$^{235}_{92}U \rightarrow ^{4}_{2}\alpha + ^{231}_{90}Th$$

Part (b)



Plutonium, ²³⁹Pu

Number of neutrons: 239 - 94 = 145
 Neutron-nucleon ratio: 145 / 239 = 0.607

Uranium, ²³⁵U

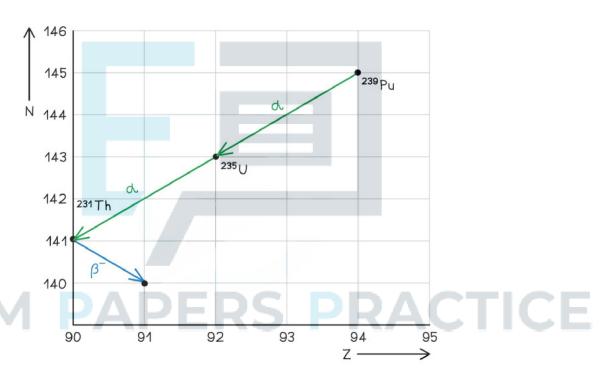
Number of neutrons: 235 - 92 = 143
 Neutron-nucleon ratio: 143 / 235 = 0.609

Thorium, ²³¹Th

Number of neutrons: 231 - 90 = 141
 Neutron-nucleon ratio: 141 / 231 = 0.610

- Thorium-231 is neutron-rich compared to uranium-235 and plutonium-239
- Therefore, it must be a β⁻ emitter

Part(c)



• The key features to draw on an N-Z graph are:

- Values for neutron number (N) on the vertical axis
- Values for proton number (Z) on the horizontal axis
- Labels for the isotopes eg. ²³⁹Pu, ²³⁵U, ²³¹Th
- o Arrows showing the direction of the decay
- Labels for the type of emission eg. α , β^-



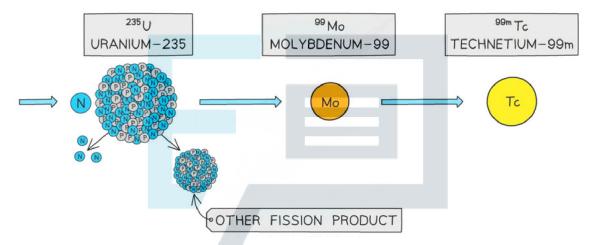
Exam Tip

Watch out for the vertical axis for the N-Z graph. Instead of N for number of neutrons this is sometimes labelled as N for **nucleon** number (total protons and neutrons) which means the decays will be represented slightly differently.

8.3.3 Nuclear Excited States

Nuclear Excited States

- In the same way that electrons can exist in excited states, nuclei can also exist in excited states
- After an unstable nucleus emits an alpha particle, beta particle or undergoes electron capture, it may emit any remaining energy in the form of a gamma photon (γ)
 - Emission of a γ photon does not change the number of protons or neutrons in the nucleus, it only allows the nucleus to lose energy
- This happens when a daughter nucleus is in an excited state after a decay
- This excited state is usually very **short-lived**, and the nucleus quickly moves to its **ground state**, either directly or via one or more lower-energy excited states



- One common application of this is the use of technetium-99m as a γ source in medical diagnosis
 - The 'm' stands for **metastable** which means the nucleus exists in a particularly stable excited state
- Technetium-99m is the decay product of molybdenum-99, which can be found as a product in nuclear reactors
- The decay of molybdenum-99 is shown below:

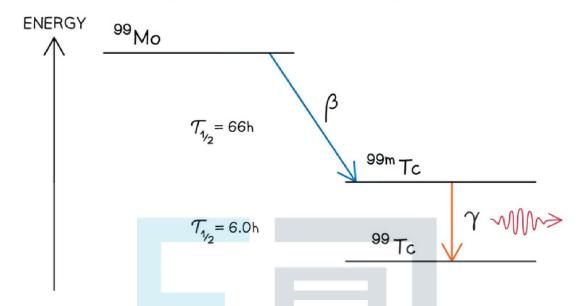
$$^{99}_{42} Mo
ightarrow \, ^{99m}_{43} Tc + ^{0}_{-1} \beta + \overline{\nu_e}$$
 $^{99m}_{43} Tc
ightarrow \, ^{99}_{43} Tc + \gamma$

- The half-life of molybdenum-99 is 66 hours
 - This is long enough for the sample to be transported to hospitals
 - Subsequently, the technetium-99m can be separated at the hospital
- Technetium-99m has a short half-life of 6 hours
 - o This is an adequate timeframe for examining a patient
 - o Plus, it is short enough to minimise damage to the patient



Nuclear Energy Level Diagrams

- Nuclear energy levels are similar to electron energy levels
- The nuclear energy level diagram of molybdenum-99 can be represented as follows:



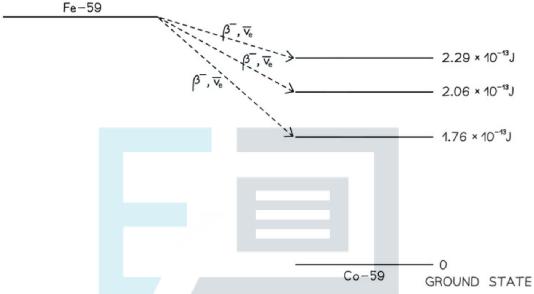
- The decay mode (usually alpha or beta) is represented by a diagonal line
- The excited state, or states, are generally stacked in descending energy order to the right of the decay





Worked Example

A nucleus of iron Fe-59 decays into a stable nucleus of cobalt Co-59. It decays by β^- emission followed by the emission of γ -radiation as the Co-59 nucleus deexcites into its ground state. The total energy released when the Fe-59 nucleus decays is 2.52×10^{-13} J. The Fe-59 nucleus can decay to one of three excited states of the cobalt-59 nucleus as shown below. The energies of the excited states are shown relative to the ground state.



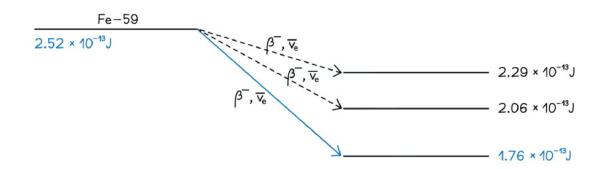
Following the production of excited states of Co-59, γ -radiation of discrete wavelengths is emitted.

- (a) Calculate the maximum possible kinetic energy of the $\beta-$ particle emitted in MeV
- (b) State the maximum number of discrete wavelengths that could be emitted
- (c) Calculate the longest wavelength of the emitted γ -radiation

Part (a)

Step 1: Identify the beta emission with the largest energy gap





Co-59 GROUND STATE

Step 2: Calculate the energy difference

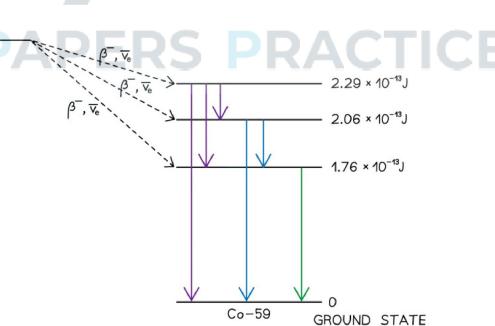
$$\Delta E = (2.52 - 1.76) \times 10^{-13} = 7.6 \times 10^{-14} \text{ J}$$

Step 3: Convert from J to MeV

$$\Delta E = \frac{7.6 \times 10^{-14}}{1.6 \times 10^{-13}} = 0.475 = 0.48 \text{ MeV}$$

 \circ 1MeV = 1.6 × 10⁻¹³ J

Part (b)

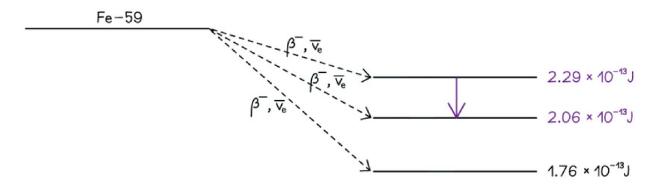


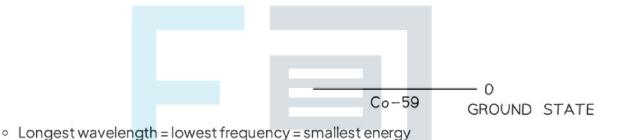
o There are 6 possible transitions, hence 6 discrete wavelengths could be emitted



Part (c)

Step 1: Identify the emission with the longest wavelength / smallest energy gap





Step 2: Calculate the energy difference

$$\Delta E = (2.29 - 2.06) \times 10^{-13} = 2.3 \times 10^{-14} \text{ J}$$

Step 3: Write down de Broglie's wavelength equation

EXAM PAPERS PRACTICE

Where:

- ∘ h = Planck's constant
- o c=speed of light

Step 4: Calculate the wavelength associated with the energy change

$$\lambda = \frac{hc}{E} = \frac{\left(6.63 \times 10^{-34}\right) \times \left(3.0 \times 10^{8}\right)}{2.3 \times 10^{-14}} = 8.6 \times 10^{-12} \text{ m}$$



8.3.4 Nuclear Radius

Estimating Nuclear Radius

Closest Approach Method

- In the Rutherford scattering experiment, alpha particles are fired at a thin gold foil
- · Some of the alpha particles are found to come straight back from the gold foil
- This indicates that there is **electrostatic repulsion** between the alpha particles and the gold nucleus

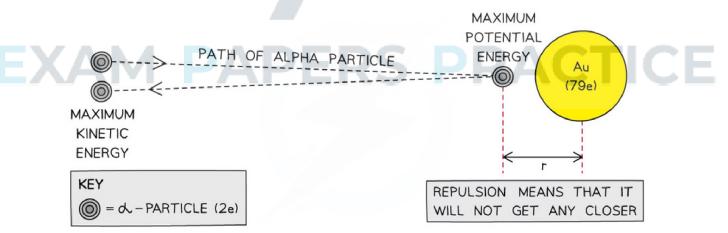
$$E_k = eV = \frac{1}{2} mv^2$$

- At the point of closest approach, r, the repulsive force reduces the speed of the alpha particles to zero momentarily
- At this point, the initial kinetic energy of an alpha particle, E_k , is **equal** to electric potential energy, E_D
- The radius of the closest approach can be found be equating the initial kinetic energy to the electric potential energy

$$E_p = \frac{Qq}{4\pi\varepsilon_0 r}$$

Equating the two equations gives:

$$E_k = E_p = \frac{1}{2} m v^2 = \frac{Qq}{4\pi\epsilon_0 r}$$



Pros & Cons of Closest Approach Method

Advantages

- Alpha scattering gives a good estimate of the upper limit for a nuclear radius
- The mathematics behind this approach are very simple
- The alpha particles are scattered only by the protons and not all the nucleons that make up the nucleus



Disadvantages

- This method does not give an accurate value for nuclear radius as it will always be an overestimate
 - This is because it measures the nearest distance the alpha particle can get to the gold nucleus, not the radius of it
- Alpha particles are hadrons, therefore, when they get close to the nucleus they are affected by the strong nuclear force and the mathematics do not account for this
- The gold nucleus will recoil as the alpha particle approaches
- Alpha particles have a finite size whereas electrons can be treated as a point mass
- It is difficult to obtain alpha particles which rebound at exactly 180°
 - o In order to do this, a small collision region is required
- The alpha particles in the beam must all have the exact same initial kinetic energy
- The sample must be extremely thin to prevent multiple scattering

Electron Diffraction Method

• Electrons accelerated to close to the speed of light have wave-like properties such as the ability to diffract and have a de Broglie wavelength equal to:

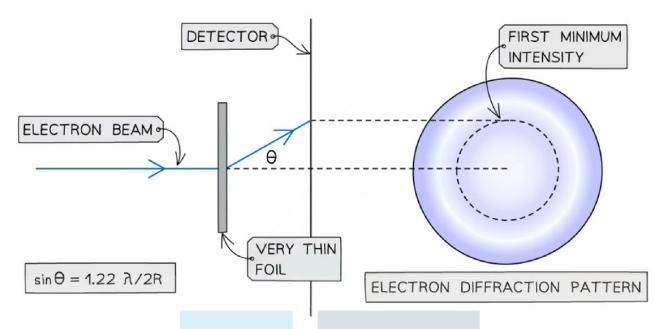
$$\lambda = \frac{h}{mv}$$

- · Where:
 - h = Planck's constant
 - m = mass of an electron (kg)
 - $v = \text{speed of the electrons (m s}^{-1})$
- The diffraction pattern forms a central bright spot with dimmer concentric circles around it
- From this pattern, a graph of intensity against diffraction angle can be used to find the diffraction angle of the first minimum
- Using this, the size of the atomic nucleus, R, can be determined from:

$$\sin \theta = \frac{1.22\lambda}{2R}$$

- Where:
 - \circ θ = angle of the first minimum (degrees)
 - $\lambda = \text{de Broglie wavelength (m)}$
 - R = radius of the nucleus (m)





Pros & Cons of Electron Diffraction Method

Advantages

- Electron diffraction is much more accurate than the closest approach method
- This method gives a direct measurement of the radius of a nucleus
- Electrons are leptons; therefore, they will not interact with nucleons in the nucleus through the strong nuclear force as an alpha particle would

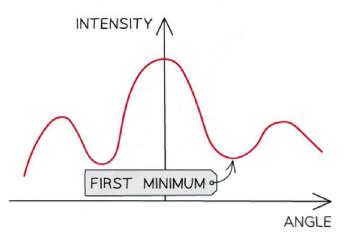
Disadvantages

- Electrons must be accelerated to very high speeds to minimise the de Broglie wavelength and increase resolution
 - This is because significant diffraction takes place when the electron wavelength is similar in size to the nuclear diameter
- Electrons can be scattered by both protons and neutrons
 - If there is an excessive amount of scattering, then the first minimum of the electron diffraction can be difficult to determine



Electron Diffraction by a Nucleus

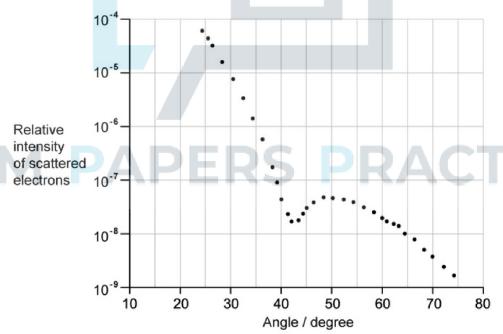
• The graph of intensity against angle obtained through electron diffraction is as follows:





Worked Example

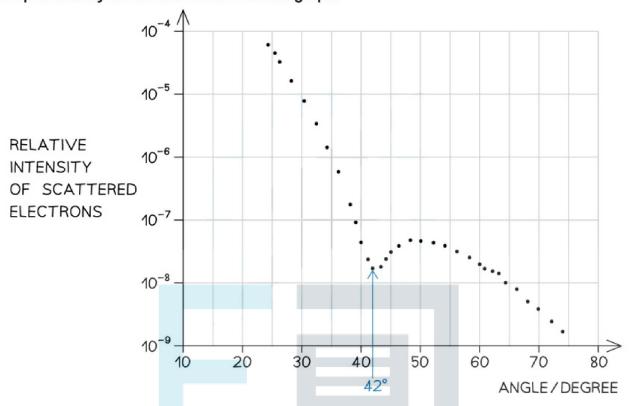
The graph shows how the relative intensity of the scattered electrons varies with angle due to diffraction by the oxygen-16 nuclei. The angle is measured from the original direction of the beam.



The de Broglie wavelength λ of each electron in the beam is 3.35×10^{-15} m. Calculate the radius of an oxygen-16 nucleus using information from the graph.



Step 1: Identify the first minimum from the graph



• Angle of first minimum, $\theta = 42^{\circ}$

Step 2: Write out the equation relating the angle, wavelength, and nuclear radius

$$\sin \theta = \frac{0.61\lambda}{R}$$

Step 3: Calculate the nuclear radius, R

$$R = \frac{0.61\lambda}{\sin \theta} = \frac{0.61 \times (3.35 \times 10^{-15})}{\sin 42^{\circ}} = 3.1 \times 10^{-15} \,\mathrm{m}$$



8.3.5 Closest Approach Estimate

Closest Approach Estimate

- The Coulomb equation can be used to give an estimate for the radius of a nucleus
- When the alpha particle reaches its closest approach (to the nucleus), all of its kinetic energy has been converted to electric potential energy
- Initially, the alpha particles have kinetic energy equal to:

$$E_k = eV = \frac{1}{2} \, mv^2$$

• The electric potential energy between the two charges is equal to:

$$E_p = \frac{Qq}{4\pi\varepsilon_0 r}$$

• At this point, the kinetic energy E_k lost by the α particle approaching the nucleus is **equal** to the potential energy gain E_p , so:

$$E_k = E_p = \frac{1}{2} m v^2 = \frac{Qq}{4\pi\varepsilon_0 r}$$

- · Where:
 - o m = mass of an α particle (kg)
 - $v = initial speed of the \alpha particles (m s^{-1})$
 - q = charge of an α particle (C)
 - Q = charge of the nucleus being investigated (C)
 - r = the radius of closest approach (m)
 - ε₀ = permittivity of free space



Worked Example

The first artificially produced isotope, phosphorus-30 ($_{15}$ P) was formed by bombarding an aluminium-27 isotope ($_{13}$ Al) with an α particle.

For the reaction to take place, the α particle must come within a distance, r, from the centre of the aluminium nucleus.

Calculate the distance, r, if the nuclear reaction occurs when the α particle is accelerated to a speed of at least 2.55×10^7 m s⁻¹.

Step 1: List the known quantities

- Mass of an α particle, $m = 4u = 4 \times (1.66 \times 10^{-27})$ kg
- Speed of the α particle, $v = 2.55 \times 10^7$ m s⁻¹
- Charge of an α particle, $q = 2e = 2 \times (1.6 \times 10^{-19})$ C
- Charge of an aluminium nucleus, $Q = 13e = 13 \times (1.6 \times 10^{-19}) C$
- Permittivity of free space, $\varepsilon_0 = 8.85 \times 10^{-12} \, \text{F m}^{-1}$



Step 2: Write down the equations for kinetic energy and electric potential energy

$$\frac{1}{2}mv^2 = \frac{Qq}{4\pi\varepsilon_0 r}$$

Step 3: Rearrange for distance, r

$$r = \frac{2Qq}{4\pi\varepsilon_0 m v^2}$$

Step 4: Calculate the distance, r

$$r = \frac{2 \times 13 \times \left(1.6 \times 10^{-19}\right) \times 2 \times \left(1.6 \times 10^{-19}\right)}{4\pi \times (8.85 \times 10^{-12}) \times 4 \times (1.66 \times 10^{-27}) \times (2.55 \times 10^{7})^{2}}$$

$$r = 2.77 \times 10^{-15} \text{ m}$$



Exam Tip

Make sure you're comfortable with the calculations involved with the alpha particle closest approach method, as this is a common exam question. You will be expected to remember that the charge of an α is the charge of 2 protons (2 x the charge of an electron)



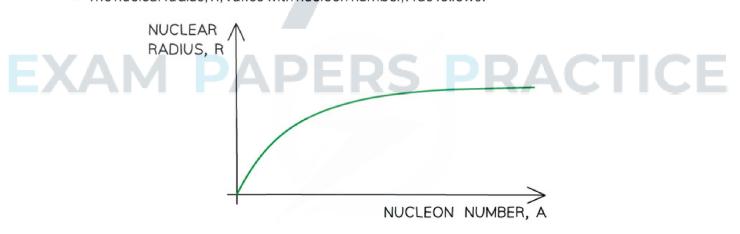
8.3.6 Nuclear Radius Equation

Nuclear Radius Values

• The radius of some nuclei are shown in the table below:

Element	Nuclear radius R / 10 ⁻¹⁵ m	Mass number A
Carbon	2.66	12
Silicon	3.43	28
Iron	4.35	56
Tin	5.49	120
Lead	6.66	208

- In general, nuclear radii are of the order 10⁻¹⁵ m or 1 fm
- The nuclear radius, R, varies with nucleon number, A as follows:



• The key features of this graph are:

- o The graph starts with a steep gradient at the origin
- Then the gradient gradually decreases to almost horizontal
- · This means that
 - o As more nucleons are added to a nucleus, the nucleus gets bigger
 - However, the number of nucleons A is **not** proportional to its size r

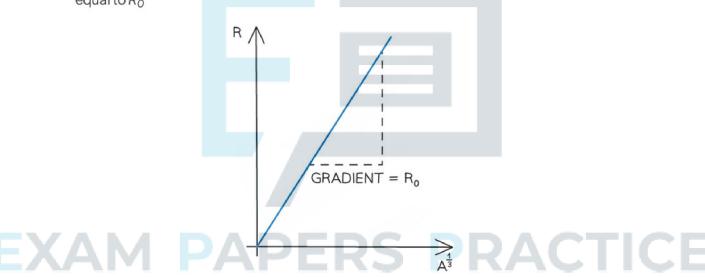


Radius v Nucleon Number

- The radius of nuclei depends on the nucleon number, A of the atom
- This makes sense because as more nucleons are added to a nucleus, more space is occupied by the nucleus, hence giving it a larger radius
- The exact relationship between the radius and nucleon number can be determined from experimental data
- By doing this, physicists were able to deduce the following relationship:

$$R = R_0 A^{\frac{1}{3}}$$

- · Where:
 - R = nuclearradius (m)
 - A = nucleon / mass number
 - \circ R₀ = constant of proportionality = 1.05 fm
- Plotting a graph of R against $A^{1/3}$ gives a straight line through the origin with the gradient equal to R_0



• It is also possible to plot a logarithmic graph of the relationship which can be derived as follows:

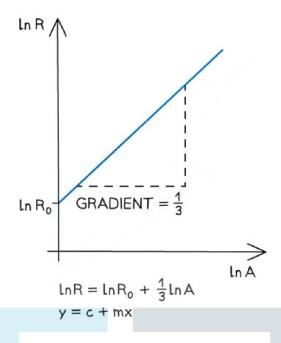
$$\ln R = \ln \left(R_0 A^{1/3} \right)$$

$$\ln R = \ln R_0 + \ln (A^{1/3})$$

$$\ln R = \ln R_0 + 1/3 \ln A$$

- Therefore, a graph of ln R against ln A yields a straight line
- Comparing this to the straight-line equation: y = mx + c
 - \circ y = lnR
 - \circ x = InA
 - o m (the gradient) = 1/3
 - \circ c (y-intercept) = $\ln R_0$





?

Worked Example

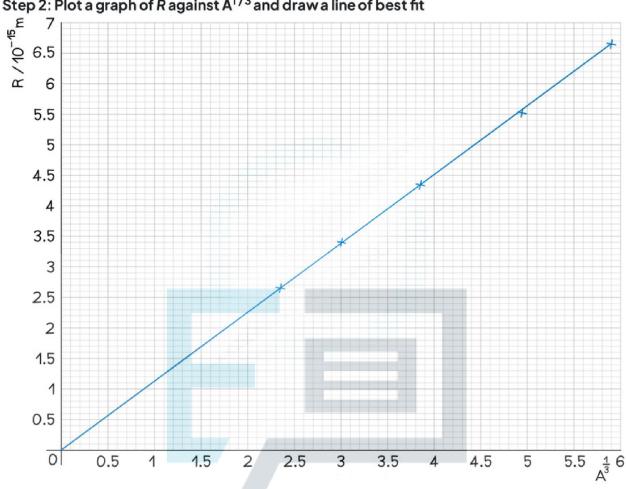
Verify the experimental relationship between R and A using the data from the table above and estimate a value of R_0 .

Step 1: Add a column to the table to determine the values for $A^{1/3}$

Element	Nuclear radius R / 10 ⁻¹⁵ m	Mass number A	Mass number A ^{1/3}
Carbon	2.66	12	2.29
Silicon	3.43	28	3.04
Iron	4.35	56	3.83
Tin	5.49	120	4.93
Lead	6.66	208	5.92

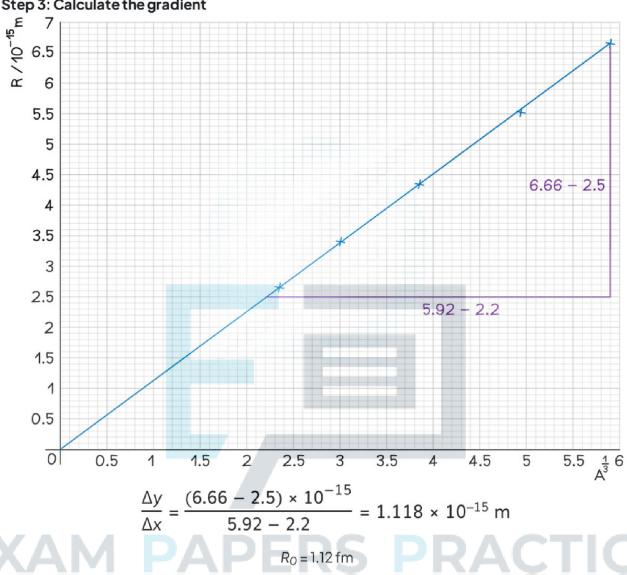


Step 2: Plot a graph of R against $A^{1/3}$ and draw a line of best fit





Step 3: Calculate the gradient



8.3.7 Nuclear Density

Constant Density of Nuclear Material

• Assuming that the nucleus is spherical, its volume is equal to:

$$V = \frac{4}{3} \pi R^3$$

• Where R is the nuclear radius, which is related to mass number, A, by the equation:

$$R = R_0 A^{\frac{1}{3}}$$

- Where R₀ is a constant of proportionality
- · Combining these equations gives:

$$V = \frac{4}{3} \pi \left(R_0 A^{\frac{1}{3}} \right)^3 = \frac{4}{3} \pi R_0^3 A$$

- Therefore, the nuclear volume, V, is proportional to the mass of the nucleus, A
- Mass (m), volume (V), and density (ρ) are related by the equation:

$$\rho = \frac{m}{V}$$

• The mass, m, of a nucleus is equal to:

$$m = Au$$

- · Where:
 - A = the mass number
 - ∘ u = atomic mass unit
- Using the equations for mass and volume, nuclear density is equal to:

$$\rho = \frac{Au}{\frac{4}{3}\pi R_0^3 A} = \frac{3u}{4\pi R_0^3}$$

- Since the mass number A cancels out, the remaining quantities in the equation are all constant
- Therefore, this shows the density of the nucleus is:
 - Constant
 - Independent of the radius
- The fact that nuclear density is constant shows that nucleons are evenly separated throughout the nucleus regardless of their size



Nuclear Density

• Using the equation derived above, the density of the nucleus can be calculated:

$$\rho = \frac{3u}{4\pi R_0^3}$$

- · Where:
 - Atomic mass unit, $u = 1.661 \times 10^{-27}$ kg
 - Constant of proportionality, $R_0 = 1.05 \times 10^{-15}$ m
- · Substituting the values gives a density of:

$$\rho = \frac{3 \times (1.661 \times 10^{-27})}{4\pi (1.05 \times 10^{-15})^3} = 3.4 \times 10^{17} \text{ kg m}^{-3}$$

- The accuracy of nuclear density depends on the accuracy of the constant R_0 , as a guide nuclear density should always be of the order 10^{17} kg m⁻³
- Nuclear density is significantly larger than atomic density, this suggests:
 - o The majority of the atom's mass is contained in the nucleus
 - The nucleus is very small compared to the atom
 - · Atoms must be predominantly empty space